# Orbits of Artificial Earth Satellites Used in the Intersputnik System with Optimum Position for Bulgaria 

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This is an examination of the problems related to the determination of the optimum elliptic orbit for Bulgaria. Graphs have been given of geostationary and elliptic orbits with different longitudes of the apogee with respect to Sofia. The authors have analysed the conditions for communication of the other participants in the Intersputnik system when the satellite operates at an optimum orbit for our country, Quantitative evaluations have been given of the conditions for communication with an artificial earth satellite on elliptic and geostationary orbits.

## Introduction

One of the forthcoming objectives of Bulgaria is the construction of an earth station (ES) for communications through artificial earth satellites (AES). In her capacity of participant in the Intersputnik international system for sate!lite communications, Bulgaria will operate with the satellites of that system and is interested in obtaining the optimum or near-optimum choice of the elliptic orbit to be used.

One basic variable parameter in the optimization of the conditions for operation with AES on an elliptic orbit is the position of the orbital plane in relation to Bulgaria. The distance between the meridian of the orbit apogee $\lambda_{A}$ and the latitude $\lambda_{E S}$ of the ES determines the proximity of the plane of the elliptic orbit.

The aim of our present work was to determine $\lambda_{A}$ in such a manner as to obtain optimum conditions for communication between the AES and the ES of Bulgaria.

The optimum elliptic orbit is the one which ensures the following:

1. Maximum time for communication performance with AES;
2. Minimum in size biological zone of the ES;
3. Minimum by-pass angle in a horizontal direction.

This results in improvement of the electromagnetic compatibility with RRL operating or intended for operation in the band of joint operaticn with ES.
4. The noise temperature introduced through the aerial of the ES station from the atmosphere should be minimal.

## Basic Dependences

The radtus $r_{0}$ of the region of possible radio-communication between an ES and an AES travelling along an elliptic orbit of the Molniya-i type is determined by the dependence (Fig. 1)

$$
\begin{equation*}
r_{0}=\frac{a_{i}}{180^{0}} \pi R, \tag{1}
\end{equation*}
$$

whete $2 \alpha_{i}$ is an arc angle of the radiovisibility region from the satellite;
$i=1,2, \ldots, n$-points from the elliptic orbit; and $R$ is the average Earth radius ( $6,370 \mathrm{~km}$ ).

The angle $\gamma$ chatacterizes the range of vision from an $A E S$

$$
\begin{equation*}
\gamma_{t}=\arcsin \frac{R \cos \frac{\beta_{i}}{R+\Pi} \frac{H}{s a t}}{} \operatorname{lgrad} \tag{2}
\end{equation*}
$$

$$
\begin{equation*}
\gamma_{t} \leq \theta_{6.5 p}^{\circ} \tag{3}
\end{equation*}
$$

where $2 \gamma_{i}$ is the span of the radiovisibility apex angle from an AES upon its travel along an elIftic orbit.
$2 \Theta_{0.5 \mathrm{p}}^{1}$ is the width of the diagram of $A E S$ antentia ofiented for operation at a hali-power level. According to [1], $200_{0.5}^{\circ} \rho_{\mathrm{p}}=20^{\circ}$.


Fig. 1

$$
\begin{equation*}
H_{\mathrm{sat}}=r_{t}-R \tag{4}
\end{equation*}
$$

is the height of the satelite above the Earth's surface; $r_{i}$ - radius vector of an $i$ point of the elliptie orbit where the satellite is to be found at the particular moment; and $\beta_{i}$ is the minimum angle of operation of the aerial of the ES above the horizots.

In view of considerations for reducing the noise temperature of the aerial, as introduced from the Earth, $\beta_{i} \geqslant 5^{\circ}$.

The dependence between the above angles is determined from

$$
\begin{equation*}
a_{l}=90^{\circ}\left(\gamma_{t}+\beta_{i}\right)[g r a d] \tag{5}
\end{equation*}
$$

and

$$
\begin{equation*}
\beta_{6}=\arccos \frac{R+H^{s a t}}{R} \sin \gamma[\mathrm{grad}] . \tag{6}
\end{equation*}
$$

## Determining the Optimum Elliptic

Orbit for Bulgaria
The town of Sofia (geographic coordinates $\lambda=23^{\circ} \mathrm{e} .1$. and $\varphi=43^{\circ}$ n. 1.) was-selected as the observation point in determining the visibility of the satellite pass along a certain elliptic orbit. The position of the satelite in a vertical plane is
determined by the angle above the horizon $\Delta^{\circ}$, while the direction toward the hotizon is determined by the azimuth angle $\varepsilon$.

The angle $\Delta^{\circ}$ is determined by the dependence

$$
\begin{equation*}
\Delta=\operatorname{arctg} \frac{A M}{\mathrm{CD}+\lambda \mathrm{D}}=\operatorname{arctg} \frac{\cos \theta-R i r}{\sin \theta}[\operatorname{lgrad}] \tag{7}
\end{equation*}
$$

where $r$ is the radius-vector of the point at which the AES is to be found, and

$1 \cdot 18.2$ $\theta$ is the geocentric angle between the point of observation $C(\lambda, \varphi)$ and the projection of the satellite on the Earth's surface $N\left(\lambda_{\text {sat }}\right.$, $\varphi_{\text {sat }}$ ). The angle $\Theta$ determines the distance between the points $C$ and $N$ by the dependence

$$
\begin{equation*}
Q-\frac{\theta}{180^{\prime \prime}} \cdot \pi R . \tag{8}
\end{equation*}
$$

The angle $\Theta$ is determined from the spherical triangle NCP (Fig. 2)
(9)

$$
\Theta=\arccos \left[\sin \varphi \cdot \sin \varphi_{\mathrm{sat}}+\cos \varphi \cdot \cos \varphi_{s, t} \cdot \cos \Delta \lambda\right][\operatorname{grad}],
$$

where $\lambda$ and $\varphi$ are the coordinates of the observation point;
$\lambda_{\text {sat }}$ and $\varphi_{\text {sat }}$ are geographical coordinates of the satelitite projection:

$$
\begin{equation*}
\Delta \lambda=\lambda_{\text {sat }}-\lambda . \tag{10}
\end{equation*}
$$

The acimuth angle $\xi$, \{aken in a North-East - South-West direction, is determined from

$$
\begin{equation*}
\xi=\Delta A \tag{11}
\end{equation*}
$$

when $\lambda_{\text {sat }}$ is to the east of the meridian $\lambda=23^{\circ}$ e. 1 . and from

$$
\begin{equation*}
\xi=360^{\circ}-\Delta A \tag{12}
\end{equation*}
$$

when $\lambda_{\text {sat }}$ is to the West of the meridian $\lambda=23^{\circ} \mathrm{e}$. I., while $\Delta \Lambda$ is determined from the spherical triatgle $N C$ ?

$$
\begin{equation*}
A A=\arccos \frac{\sin \varphi \operatorname{sat}-\sin \varphi \cdot \cos \theta}{\sin \theta \cdot \cos \varphi} \frac{\operatorname{lgrad}] .}{} \tag{13}
\end{equation*}
$$

The geographic coordinates of the satelite $\lambda_{\text {sat }}, \varphi_{\text {sai }}$ at any moment of its movement along the elliptic orbit are determined by the geocentric projection of the orbit on the Earth's surface. Figure 3 shows the geocentric profection of the odd elliptic trajectory, according to [2], due account being taken of the Earth's movement.

The projection of the even elliptic trajectory is a continualion of the odd one and has the position of the apogee $\lambda_{A}^{\prime \prime}$ :

$$
\begin{equation*}
\lambda_{A}^{\prime \prime}=\lambda_{A}^{\prime}+180^{\circ} \tag{14}
\end{equation*}
$$

Figure 4 shows graphically presented elliptic orbits with differen longitude of the apogec $\lambda_{A}^{\prime}\left(\lambda_{A}^{\prime}\right)$, as they are seen from the selected observation point. The position of the satellite is determined in relation to the moment of time in which the ALS passes through the pertigee point $\left(t_{0}=00 \mathrm{~h}\right)$.

Table 1 gives the basic quantities which are characteristic of the elliptic orbit, namely: the radius-vector $r, V^{\prime \prime}$ the angle between the direction to the perigee and $r$ ), the radiovisibility zone from the satellite $\left(\alpha, r_{0}\right)$ and the distance from the observation point to the projection of the satellite on the Earth's surface ( $\Theta, \rho)$ for the selected elliptic orbits with different longitude of the apogee in function of absolute time.

For the purpose of determining the duration of the session for communication with the AES, we conpare the visual zone of the AES from (1) and the distance to the undersatellite point from (8) (Table 1).

1. At $p>r_{0}(\Theta>a)$ the AES cannot "see" us with ils aerials. Our visibility toward the satellite is determined by $\Delta$ from (7) and $\varepsilon$ from (11) or (12), and depends on the overlap angle to the horizon. When it is possible to ensure a minimum covering angle ( $\beta_{\text {rnin }}=$ : $5^{\text {b }}$ ) for all orbits shown on Fig. 4, we can follow the movemeni of the satellite within an approximately 11 hour sector. In order to realize the communication session it is necessary to adjust the diagram of directed opetation of the satel-


Fig. 3 tite's aerial.
2. At $\rho \leqslant r_{0}(\Theta \leqslant \alpha)$ the AES satellite can be used to establish communicalion with member-country of the Intersputnik system. The boundary line $\rho=r_{0}$ at $\beta=5^{\circ}$ and $\gamma=10^{\circ}$, plotted by a broken line on Fig. 4, determines the duration of the communication session as shown in Table 2. The results obtained show that upon using one AES travelling along an elliptic orbit with different positions of the apogee, the total time of the communication session is a sum of two variable components.

The dimensions of the biological zone of the ES for protection from the irradiation of the aerial within the microwave band for persons not professionally involved in radiation and for the population is determined, as regards intensity, at $1 \mu W / \mathrm{cm}^{2}$ [3].

The size of the biozone depends on the power of the operating transmitters and on the values of the operative angles in horizontal and vertical directions. The dimensions of the biozones for the selected orbits at transmitter power $P_{t_{r}}$ $=10 \mathrm{~kW}$, as well as the bypass angles from the aerial in a horizontal direction, are given in Table 2:

The noise temperature of the aerial $T_{\text {na }}$ (introduced from dry atmosphere) is significant to the quality of the signal received from the satellite, whose var tue is of the order of $10^{-14} \mathrm{~W} / \mathrm{m}^{2}$. It depends on the angle above the horizont $\Delta^{\circ}$ at which the aerial is operating
Table 1

| No. | $\begin{gathered} \text { Time after } \\ \text { the perigee. } \\ \text { (h; mini }] \end{gathered}$ | $V^{\circ}, 10$ | $\underset{\substack{r \\ \hline \\[\mathrm{~km}+\mathrm{R}+\mathrm{H}}}{ }$ | ${ }^{\text {[9] }}$ | ${ }_{\substack{\text { com } \\\left[\mathrm{k}_{\mathrm{om}}\right.}}$ |  |  | $\left\lvert\, \begin{gathered} { }^{2} A=20^{\circ}{ }^{\circ} \mathrm{e} . \mathrm{e} .1 \\ 160^{\circ} \mathrm{w} .1 \end{gathered}\right.,$ |  | $\left\|\begin{array}{c} \lambda_{\mathrm{A}}=30^{\circ} \mathrm{e} . \mathrm{e} .1 . \\ 155^{\circ} \mathrm{w} .1 . \end{array}\right\|$ |  |  |  |  |  | $\left\|\begin{array}{r} \lambda_{\mathrm{A}}=82^{\circ} \text { e. } 1 . \\ 98^{\circ} \text { w. } \end{array}\right\|$ |  |  |  | $\begin{gathered} \lambda_{A}=10^{\circ}=., 1.1 . \\ 70^{\circ} w, 1 . \end{gathered}$ |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  |  |  |  | $\theta$ | $\mid$ elkm] |  | $01 \mid e[\mathrm{~km}]$ | $\theta 0^{\circ} \mathrm{O}$ | m] | $\left.{ }^{100}\right]$ | e[km] |  | ] $19[\mathrm{~km}]$ | 6["] | elkm] | $8[9$ | efkm | $\theta\left[{ }^{\circ}\right]$ | [lkm] |
| 1 | 20 m | 54,5 | 8376 | 3,2 | 350 | 63,1 | 015 | 58 | 6450 | 54 | 6000 | 56 | 6230 | 59 | 6560 | 66 | 7340 | 1 | 1 |  |  |
| 2 | 30 m | 76 | 9649 | 5,3 | 580 | 45,5 | 5058 | 39 | 335 | 38 | 4225 | 39,5 | 4390 | 48 | 5340 | 61 | 6780 | 70 | 778 |  |  |
| 3 | 1 h | 117 | 18050 | 19,4 | 2153 | 26 | 2890 | 19 | 2110 | 20,5 | 2280 | 25 | 2780 | 37,5 | 3,5 4170 | 53 | 5890 | 62,5 | 6950 | 73 | 11 |
| 4 | 1 h 30 | 138 | 26573 | 36,4 | 4038 | 14 | 1560 | 5 | 556 | 10 | 1110 | 16,5 | 1840 | 31,5 | 53500 | 47 | 5230 | 56 | 6230 | 67 | 7450 |
| 5 | 2 h | 149, | 33140 | 59,5 | 6615 | 15 | 1670 | 6 | 667 | 9 | 1000 | 15 | 1670 | 28,5 | 5.3170 | 43 | 4780 | 51 | 567 | 59,5 | 6615 |
| 6 | 3 h | 160, | 39387 | 76 | 8460 | 20 | 2220 | 14 | 1560 | 15 | 1680 | 18,5 | 2060 | 28,5 | 53170 | 40,5 | 4500 | 48 | 5340 | 55 | 5 |
| 7 | 6 h | 180 | 45961 | 77 | 856 | 25 | 2780 | 21 | 2335 | 21 | 2335 | 23 | 2560 | 29,5 | 53280 | 39 | 4340 | 45 | 501 | 51 | 5670 |
| 8 | 9 h | 200,2 | 39103 | 76 | 8460 | 21,5 | 2390 | 14,5 | 4,5 1610 | 14,5 | 1610 | 17 | 1890 | 26 | 2880 | 38 | 4225 | 43 | 4780 | 53 | 5890 |
| 9 | 10 h | 210 | 33324 | 59,5 | 6615 | 15,5 | 1720 | 8 | 880 | 7,3 | 810 | 11,5 | 1280 | 24 | 2670 | 39 | 4340 | 47 | 523 | 56 | 6225 |
| 10 | 10h30 | 221,5 | 26811 | 37 | 4100 | 21 | 2330 | 8 | 890 | 5 | 556 | 10 | 1110 | 24,5 | . 2720 | 40 | 4450 | 49 | 5460 | 59 | 6560 |
| 11 | 11 h | 241 | 18680 | 20,5 | 2280 | 28,5 | 3170 | 20 | 2220 | 19,7 | 2190 | 27 | 2450 | 35 | 3890 | 50 | 5560 | 60 | 668 | 70 | 7780 |
| 12 | $11 \mathrm{~h} \mathrm{30m}$ | 280 | 10615 | 6,1 | 750 | 41,5 | 4610 | 38 | 4225 | 39,5 | 4390 | 43 | 4780 | 53 | 5900 | 68 | 7560 | 71 | 791 |  |  |
| 13 | 14 h | 149,5 | 33140 | 59,5 | 6615 | 1 | 1 | 1 | 1 | 1 | 1 | 1 | 1 | 1 | 1 | 74 | 6230 | 68 | 7560 | 59,5 | 6615 |
| 14 | 15 h | 160,5 | 39387 | 76 | 8150 | 79 | 8780 | 80,5 | ,5 8950 | ${ }^{80,3}$ | 89,30 | 79,3 | 8320 | 70, 3 | [8370 | 68 | 7560 | 63 | 7000 | 54 | 6000 |
| 15 | 18 h | 180 | 45961 | 77 | 8560 | 73 | 8115 | 74 | 8230 | 73,5 | 8170 | 73 | 8115 | 70 | 7780 | 64 | 7115 | 59,5 | 6610 | 54 | 6000 |
| 16 | 21 h | 200,2 | 39103 | 76 | 8460 | 78 | 8670 | 80,4 | ,4 8940 | 80,5 | 8950 | 80 | 8390 | 76,3 | ${ }^{8480}$ | 7. | 7780 | 64,5 | 7170 | 58 | 6450 |
| 17 | 22 h | 210 | 33324 | 59,5 | 6 | 1 | / |  |  |  |  |  |  |  |  | $\overline{77}$ |  | 71 |  | 63 |  |


Table 2

|  | $\begin{array}{\|l\|l} \text { Long tude } \\ \text { Lof thice } \\ \text { apoge }\left[{ }^{\circ}\right] \end{array}$ | Time of commmurication session: Iroun -- to (b) |  | Total time <br> (h) | Angle to the horizonand noise temp. $\left[\begin{array}{l}\text { [li } ;[k]\end{array}\right.$ |  | $\begin{gathered} \text { Rypass } \\ \text { angict } \\ 1 y_{i} \end{gathered}$ | Biozone [m! |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | 1 st tevoiution | 2nd revolution |  | 1st rev. | 2nd tev. |  | E | w | $N$ | s |  |
| 1 | $\begin{array}{\|c} 0^{\circ} \mathrm{e} .1 . \\ 180^{\circ} \mathrm{w} .1 . \end{array}$ | 1. $05 \mathrm{~mm} \div 10 \mathrm{~h} 50 \mathrm{~m}-9 \mathrm{~h} 45 \mathrm{~m}$ | 16h33m-20h06m-3b33m | 13h 18m | $65(2.8)$ | 5(28.0) | 105 | 760 | 670 | 200 | 720 | 2750 |
| 2 | $\begin{gathered} 20^{\circ} \mathrm{e} . \mathrm{l} \\ 160^{\circ} \mathrm{w} . \mathrm{i} . \end{gathered}$ | $1 \mathrm{~h} 00 \mathrm{~m} \div 1 \mathrm{~h} 00 \mathrm{~m}-10 \mathrm{~h} 00 \mathrm{~ms}$ | 17h00m $\div 19 \mathrm{~h} 25 \mathrm{~m}-2 \mathrm{~h} 15 \mathrm{~mm}$ | 12h15m | 70(2.7) | $5(28.0)$ | 325 | 720 | 720 | 1200 | 720 | 2770 |
| 3 | $\begin{aligned} & 30^{\circ} \mathrm{c}, 1 . \\ & 150^{\mathrm{o}} \mathrm{w} 1 . \end{aligned}$ | 1h01m $-11 \mathrm{~h} 00 \mathrm{~m}-9 \mathrm{~h} 59 \mathrm{~m}$ | 16h540m $-19 \mathrm{~h} 24 \mathrm{~m}-2 \mathrm{~h} 30 \mathrm{~m}$ | 12h29m | 70(2.7) | $5(28.0)$ | 151 | 720 | 720 | 1200 | 720 | 2770 |
| -4 | $\begin{gathered} 40^{\circ} \text { e. } 1 . \\ 140^{\circ} \text { w. } . ~ \end{gathered}$ | $1 \mathrm{~h} 05 \mathrm{~m} \div 10 \mathrm{~h} 54 \mathrm{mb}-9 \mathrm{~h} 49 \mathrm{~m}$ | $161132 \mathrm{~m} \div 19 \mathrm{~h} 36 \mathrm{~m}-3 \mathrm{ho3m}$ | 12h52m | 65(2.8) | 5(28.0) | 117 | 720 | 760 | 1200 | 720 | 2840 |
| 5 | $\left\lvert\, \begin{gathered} 60^{\circ} \mathrm{e} .1 \\ 120^{\circ} \mathrm{w} .1 . \end{gathered}\right.$ | 1h22m1-10h43m-9h21m | $14 \mathrm{~h} 57 \mathrm{~m} \div 20 \mathrm{~h} 45 \mathrm{~m}-5 \mathrm{h48m}$ | 15h09m | $57(3.0)$ | $8(18.0)$ | 70.5 | 700 | 800 | 1140 | 740 | 2820 |
| 6 | $\begin{aligned} & 82^{\circ} \mathrm{e} \cdot 1.1 \\ & 98^{\circ} \mathrm{w} \cdot \mathrm{l} . \end{aligned}$ | 1h48m $\div 10 \mathrm{~h} 27 \mathrm{~m}-8 \mathrm{~h} 39 \mathrm{~m}$ | $14 \mathrm{~h} 30 \mathrm{~m} \div 21 \mathrm{~h} 77 \mathrm{~m}-6 \mathrm{~h} 53 \mathrm{~m}$ | 15h32m | 453.5) | 12(12.0) | 49 | 760 | 780 | 870 | 630 | 2310 |
| 7 | $\begin{gathered} 96^{\circ} \text { e. } 1 . \\ 85^{\circ} \text { w.1. } \end{gathered}$ | 1h54m $-10 \mathrm{~h} 15 \mathrm{~mm}-8 \mathrm{~h} 21 \mathrm{~m}$ | $14 \mathrm{~h} 03 \mathrm{~m} \div 2 \mathrm{hm} 45 \mathrm{~mm}-7 \mathrm{~h} 42 \mathrm{~m}$ | 16h03m | 38(4.1) | 20(7.3) | 40 | 760 | 800 | 850 | 530 | 2150 |
| 8 | $\begin{gathered} 110^{\circ} \mathrm{e}, 1 \\ 70^{\circ} \mathrm{w} \cdot \mathrm{l} \end{gathered}$ | $2 \mathrm{ta8m} \div 9 \mathrm{~h} 56 \mathrm{~m}-7 \mathrm{h48m}$ | $14 \mathrm{~h} 00 \mathrm{ta} \div 21 \mathrm{~h} 54 \mathrm{~m}-7 \mathrm{~h} 54 \mathrm{~m}$ | 15h42m | 30(5.0) | 28(5.4) | 44 | 800 | 850 | 830 | 550 | 2280 |
| 9 | $68^{\circ}$ e. 1. | 24 | - | 24 | 23.7(6.2) |  | $\begin{gathered} \text { Perma. } \\ \text { nent } \\ 124.43^{\circ} \end{gathered}$ | 770 | 350 | 490 | 740 | 1380 |
| 10 | $10^{\circ} \mathrm{w} .1$. | 24 | - | 24 | 30.1(5.0) |  | Perma- nent 20 $223.6^{\circ}$ | 430 | 740 | 390 | 740 | 1320 |



$$
T_{\mathrm{na}}=\frac{2.5}{\sin A}[K]
$$

according to [2]. This formula is valid for the $1 \div 8 \mathrm{GHz}$ band and $\Delta>3 \div 5^{\circ}$. The values of the noise temperature of the aerial, calculated for an angle, which is the averaged value of the angles at which the aerial of the ES operates during most of the session for AES satellite communication, are plotted in Table 2.

A comparison of all optimized quantities in Table 2 shows that the elliptic orbit with $\lambda_{A}^{\prime}=-95^{\prime}$ e. l. and $\lambda_{A}^{\prime \prime}=85^{\circ} \mathrm{w}$. 1. is optimal both as regards the time of the communication session through the $A E S$ and also as regards the other parameters: the bypass angle of the actial in the horizontal plane is minimal, the area of the biozone is also minimal and the noisc temperature of the aerial $T_{\text {na }}$ is very close to the lowest calculated value.

## Geostationary Orbit

The visibility of a geostationary satellite (angles $A$ and $\xi$ ) is determined with formulae (7) to (13) for a satellite travelling along an elliptic orbit, and for a geostationary satellite $\varphi_{\text {sat }}=0$ and (9) it is as follows:

$$
\begin{equation*}
\theta=\arccos (\cos \varphi \cdot \cos \Delta \lambda][g r a d] \tag{16}
\end{equation*}
$$

At a constant height $H_{\text {sat }}=36,000 \mathrm{~km}$ and constant distance $r=H_{\text {sat }}+R$ for a geostationary satellite, (7) becomes as follows:

$$
\begin{equation*}
\Delta=\operatorname{arctg} \frac{\cos \theta-0.15}{\sin \theta}[\mathrm{grad}] \tag{17}
\end{equation*}
$$

The angle $\Delta A$ for a geostationary satellite is determined from (13) and becomes

$$
\begin{equation*}
A=\arccos \left[\frac{\operatorname{tg} \varphi}{\operatorname{tg} \theta}\right] \tag{18}
\end{equation*}
$$

The azimuth angle $\xi$ is determined from (11) and (12). Figure 5 shows a graphically presented geostationary orbit with visibitity from the selected observation point of Sofia. Broken lines show the intervals of possible position of a geostationary AES in the Intersputnik sysem from 6 to $28^{\circ} \mathrm{W}$. . . and from 68 to $95^{\circ} \mathrm{e} .1$. [4]. The positions have been designated of a western satelite at $10^{\circ} \mathrm{w}$. I. and an eastern satellite at $68^{\circ}$ e. I. which have been established as fully satisfying the needs of Bulgaria for communications over an AES with all countries in the world. For the purpose of comparison, Table 2 contains the values corresponding to the optimizing quantities.

## Analysis of the Results Obtained

The visibility of the optimum for our couniry elliptic orbit from several points with changing latitudinal parameter - Moscow (56 n. I.; $38^{\circ}$ e. 1.), Novosibirsk ( $55^{\circ}$ n. $1 . ; 82^{\circ}$ e. 1.), Warsaw ( $52^{\circ} \mathrm{n} .1 . ; 20^{\circ}$ e. 1.), Ulan Bator ( $48^{\circ} \mathrm{n} .1 . ;$ $107^{\circ}$ e. 1.), Soffa ( $43^{\circ}$ n. 1.; $23^{\circ}$ e. 1.), and Cuba ( $22^{\circ}$ n. l.; $80^{\circ}$ w. 1.) calculated according to formulae (7) to (13), is shown on Fig. 6. Given below are the calculated times for the communication session.

The conclusion may be drawn from the above data and from Fig. 6 that the conditions of operation offered by this orbit are not unfavourable both as


Fig. 6

Communication time

| Town | 1strevolation | 2nd revolution | Total |
| :---: | :---: | :---: | :---: |
| Moscow | 8 h 58 min | 7 h 38 min | 16 h 36 min |
| Novosibirsk | 9 t 34 min | 7 l | 16 h 34 mln |
| Warsaw | 8 h 38 min | 7 th 53 min | 16 h 26 min |
| Ulan Bator | 9 b 50 mia | 6 h 12 min | 16 h 02 min |
| Solia | 8 h 21 min | 7 6 42 min | 16 h 03 min |
| Cuba | -_ | 10 n 36 min | 10 h 36 min |

regards the duration of the communication session and as regards the operating argle of the aerial. The data available warant the conclusion that shifting to the south results in a decrease in the period of the session.

The determination of an elliptical orbit which is optimal for the countries belonging to the Intersputnik system has rot been the object of the present work.
lt can be seen from Fig. 5 that the ES in Sofia can establish communication with geostationary satellites at a minimum overlap angle toward the horizon of $5^{\circ}$, above the meridians from $94.3^{\circ}$ e. 1. to $48.3^{\circ}$ w. I. which are visible in a horizontal plane in the directions from $103^{\circ}$ to $257^{\circ}$. For the purpose of comparing the changes in the conditions of operation with the above Western and Eastern geostationary satellites from two points of different latitude, calculations have been made for Sofia and Warsaw.


It is possible to conclude from the comparative data presented above that, practically, the conditions of operation change only insignificantly with the change in latitude.

## Conclusion

The analysis of the results obtained shows that the optimum elliptic orbit for Bulgaria is that with apogee of an even revolution over meridian $95^{\circ} \mathrm{e} .1$. and of the odd one over meridian $85^{\circ}$ w. I.

For our latitude the geostationary satellites examined offer definitely better conditions for communication through AES.

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Оптимальные для НРБ орбиты ИСЗ, использованные в системе „Интерспутіик"

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## (Pes P м e)

Рассмотрены вопросы, связанные с определением оптимаяьной для НРБ эллиптической орбиты. Графически представлены геостационарная и эллиптические орбиты с изменением географической долготы й апогея по отғопению к Софии. Анализированы условия связи с другими участниками в системе „Интерспутник" в условиях работы со спутником на оптимальной для нашей страны орбите. Даны количественные оценки условий связи со опутниками на эллиптических и геостационарной орбитах.

